

3.1-d: Hydrogen Spectra

- Later, we will learn about selection rules for transition. For the principal quantum number, n , Δn can be any positive (absorption) or negative (for emission) integer, resulting in the very rich form of the H atom spectrum. The Rydberg formula

$$\omega = R_H \cdot \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \quad \text{where } R_H = \frac{2\pi^2 \mu_e e^4}{(4\pi\epsilon_0)^2 h^3 c} = 109,737 \text{ cm}^{-1} \quad \text{and } n_1 < n_2$$

- Emission lines fall in characteristic spectral regions

The Lyman series ($n_1 = 1$) in the ultraviolet

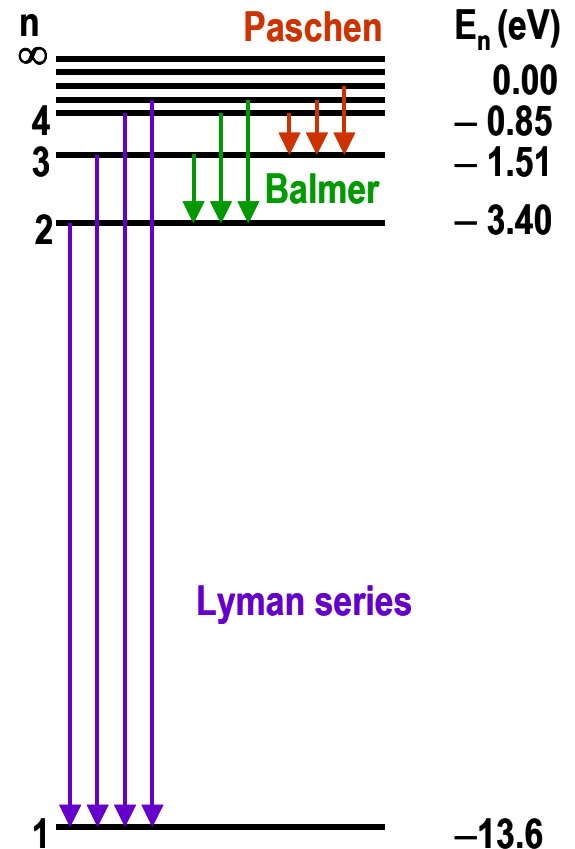
The Balmer series ($n_1 = 2$) in the visible

The Paschen series ($n_1 = 3$) in the near IR

The Balmer series



- Absorption frequencies coincide with those for emission
- Transitions having $n_1 \neq 1$ are observed only after preparation of the excited state by some means, such as electrical discharge.
- The ionization limit in absorption corresponds to a final quantum number $n_2 = \infty$. From the ground state, $I_o = R_H$, or 13.6 eV



3.1-e: The H-Atom Orbitals

$$\Psi_{n\ell m_\ell}(r, \theta, \varphi) = R_{n\ell}(r) \cdot Y_{\ell m_\ell}(\theta, \varphi)$$

Imaginary functions

$$\psi_{100} = \frac{1}{\sqrt{\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot e^{-r/a_0}$$

$$\psi_{200} = \frac{1}{4\sqrt{2\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \left(2 - \frac{r}{a_0} \right) \cdot e^{-r/2a_0}$$

$$\psi_{210} = \frac{1}{4\sqrt{2\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \frac{r}{a_0} \cdot e^{-r/2a_0} \cdot \cos\theta$$

$$\psi_{21\pm 1} = \frac{1}{8\sqrt{2\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \frac{r}{a_0} \cdot e^{-r/2a_0} \cdot \sin\theta \cdot e^{\pm i\varphi}$$

$$\psi_{300} = \frac{1}{81\sqrt{3\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \left(27 - 18\frac{r}{a_0} + 2\frac{r^2}{a_0^2} \right) \cdot e^{-r/3a_0}$$

$$\psi_{310} = \frac{1}{81} \left(\frac{2}{\pi} \right)^{1/2} \left(\frac{1}{a_0} \right)^{3/2} \cdot \left(6\frac{r}{a_0} - \frac{r^2}{a_0^2} \right) \cdot e^{-r/3a_0} \cdot \cos\theta$$

$$\psi_{31\pm 1} = \frac{1}{81\sqrt{\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \left(6\frac{r}{a_0} - \frac{r^2}{a_0^2} \right) \cdot e^{-r/3a_0} \cdot \sin\theta \cdot e^{\pm 2i\varphi}$$

$$\psi_{320} = \frac{1}{81\sqrt{6\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \frac{r^2}{a_0^2} \cdot e^{-r/3a_0} \cdot (3\cos^2\theta - 1)$$

$$\psi_{32\pm 1} = \frac{1}{81\sqrt{\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \frac{r^2}{a_0^2} \cdot e^{-r/3a_0} \cdot \sin\theta \cdot \cos\theta \cdot e^{\pm i\varphi}$$

$$\psi_{32\pm 2} = \frac{1}{162\sqrt{\pi}} \left(\frac{1}{a_0} \right)^{3/2} \cdot \frac{r^2}{a_0^2} \cdot e^{-r/3a_0} \cdot \sin^2\theta \cdot e^{\pm 2i\varphi}$$

$$\Psi_{n\ell m_\ell}(r, \theta, \varphi) = R_{n\ell}(r) \cdot Y_{\ell m_\ell}(\theta, \varphi)$$

Real functions

$$\psi_{2p_x}(r, \theta, \varphi) = \frac{1}{4\sqrt{2\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \frac{r}{a_o} \cdot e^{-r/2a_o} \cdot \sin\theta \cdot \cos\varphi$$

$$\psi_{2p_y}(r, \theta, \varphi) = \frac{1}{4\sqrt{2\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \frac{r}{a_o} \cdot e^{-r/2a_o} \cdot \sin\theta \cdot \sin\varphi$$

$$\psi_{2p_z}(r, \theta, \varphi) = \frac{1}{4\sqrt{2\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \frac{r}{a_o} \cdot e^{-r/2a_o} \cdot \cos\theta$$

$$\psi_{3p_x}(r, \theta, \varphi) = \frac{\sqrt{2}}{81\sqrt{\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \left(6\frac{r}{a_o} - \frac{r^2}{a_o^2}\right) \cdot e^{-r/3a_o} \cdot \sin\theta \cdot \cos\varphi$$

$$\psi_{3p_y}(r, \theta, \varphi) = \frac{\sqrt{2}}{81\sqrt{\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \left(6\frac{r}{a_o} - \frac{r^2}{a_o^2}\right) \cdot e^{-r/3a_o} \cdot \sin\theta \cdot \sin\varphi$$

$$\psi_{3p_z}(r, \theta, \varphi) = \frac{\sqrt{2}}{81\sqrt{\pi}} \left(\frac{1}{a_o}\right)^{3/2} \cdot \left(6\frac{r}{a_o} - \frac{r^2}{a_o^2}\right) \cdot e^{-r/3a_o} \cdot \cos\theta$$

$$\Psi_{n\ell m_\ell}(\mathbf{r}, \theta, \varphi) = R_{n\ell}(\mathbf{r}) \cdot Y_{\ell m_\ell}(\theta, \varphi)$$

Real functions

$$\psi_{3d_{z^2}}(r, \theta, \varphi) = \frac{1}{81\sqrt{6\pi}} \left(\frac{1}{a_o}\right)^{3/2} \frac{r^2}{a_o^2} \cdot e^{-r/3a_o} \cdot (3\cos^2\theta - 1)$$

$$\psi_{3d_{xz}}(r, \theta, \varphi) = \frac{\sqrt{2}}{81\sqrt{\pi}} \left(\frac{1}{a_o}\right)^{3/2} \frac{r^2}{a_o^2} \cdot e^{-r/3a_o} \cdot \sin\theta \cdot \cos\theta \cdot \cos\varphi$$

$$\psi_{3d_{yz}}(r, \theta, \varphi) = \frac{1}{81\sqrt{\pi}} \left(\frac{1}{a_o}\right)^{3/2} \frac{r^2}{a_o^2} \cdot e^{-r/3a_o} \cdot \sin\theta \cdot \cos\theta \cdot \sin\varphi$$

$$\psi_{3d_{x^2-y^2}}(r, \theta, \varphi) = \frac{1}{81\sqrt{2\pi}} \left(\frac{1}{a_o}\right)^{3/2} \frac{r^2}{a_o^2} \cdot e^{-r/3a_o} \cdot \sin^2\theta \cdot \cos 2\varphi$$

$$\psi_{3d_{xy}}(r, \theta, \varphi) = \frac{1}{81\sqrt{2\pi}} \left(\frac{1}{a_o}\right)^{3/2} \frac{r^2}{a_o^2} \cdot e^{-r/3a_o} \cdot \sin^2\theta \cdot \sin 2\varphi$$

•The wave functions

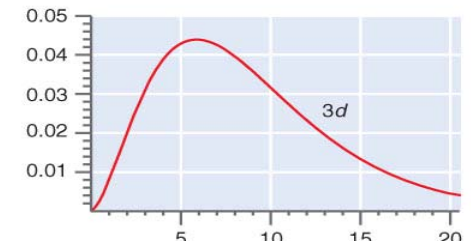
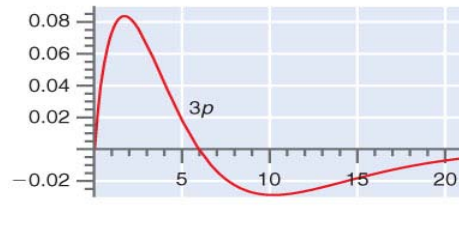
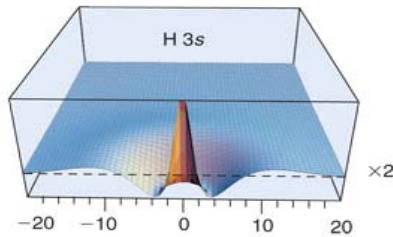
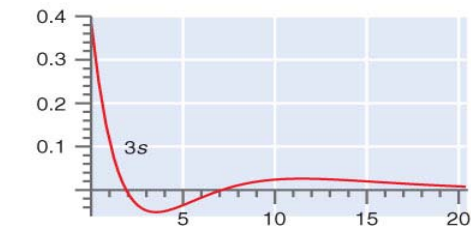
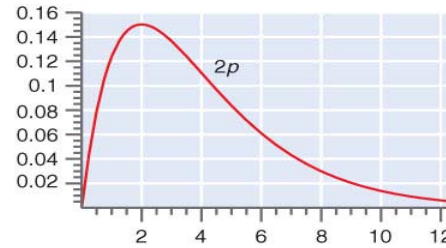
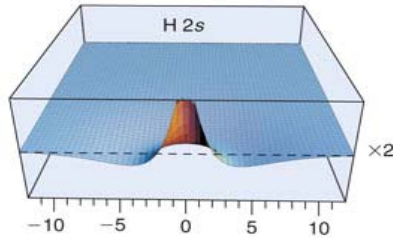
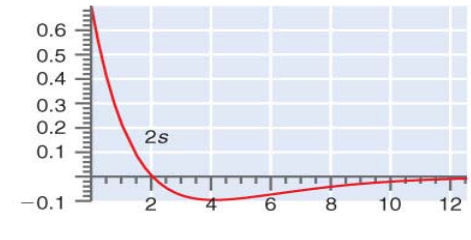
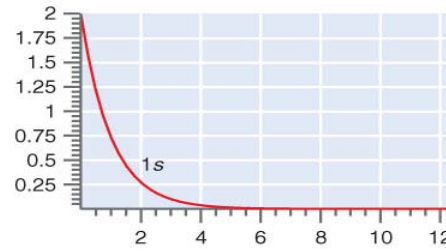
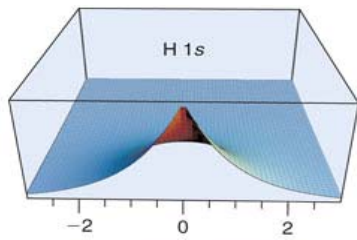
$$\Psi_{n\ell m_\ell m_s}(\mathbf{r}, \theta, \varphi, \delta) = R_{n\ell}(\mathbf{r}) \cdot Y_{\ell m_\ell}(\theta, \varphi) \cdot \begin{Bmatrix} \alpha \\ \beta \end{Bmatrix}$$

where $R_{n,\ell}(\mathbf{r})$ = the radial wave function

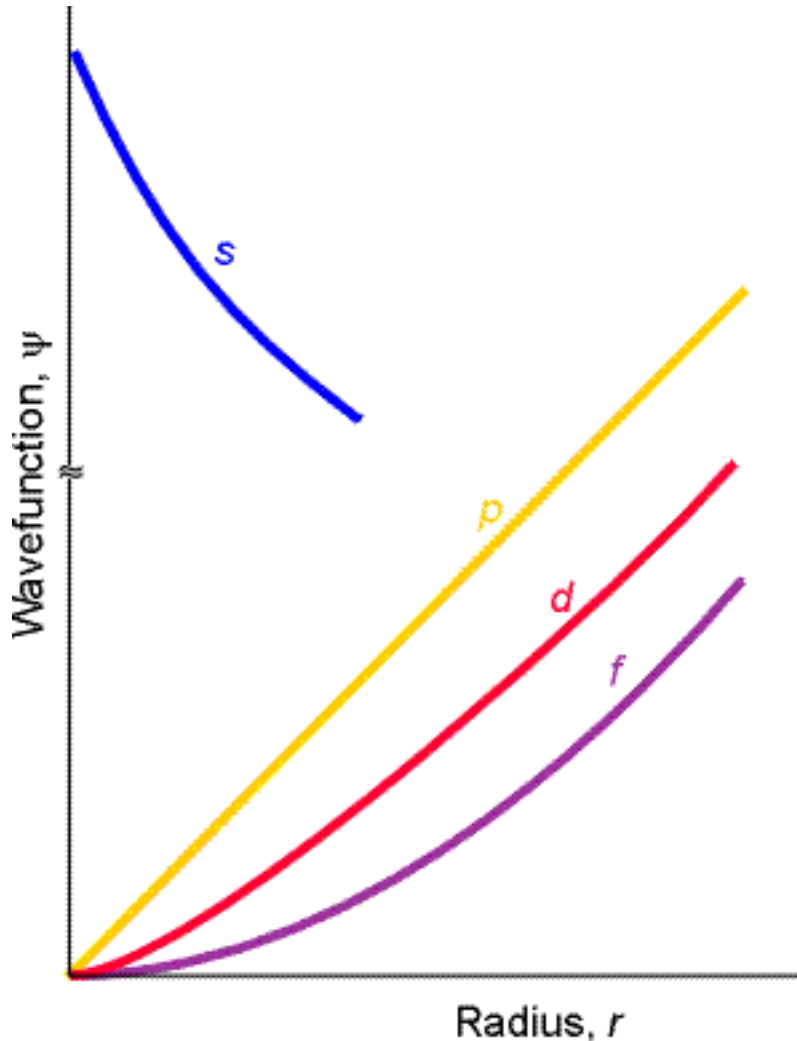
and $Y_{\ell, m_\ell}(\theta, \varphi)$ = the angular wave function

$$Y_{\ell, m_\ell}(\theta, \varphi) \propto \Theta_{\ell, m_\ell}(\cos\theta) \cdot e^{im_\ell\varphi}$$

The φ -dependence is important in giving rise to selection rules on changes in the quantum number m_ℓ



Behaviour of the orbitals near the nucleus



Close to the nucleus,

p orbitals are proportional to r

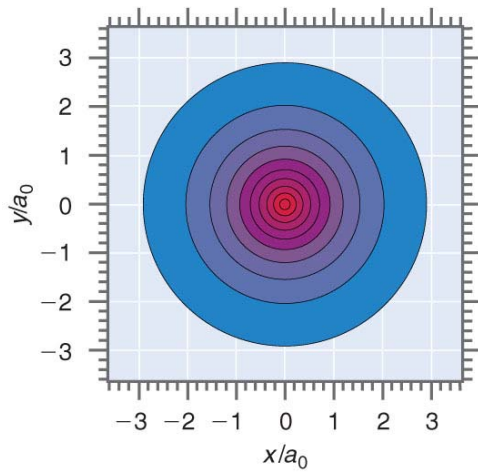
d orbitals are proportional to r^2

f orbitals are proportional to r^3

Electrons are progressively excluded from the neighborhood of the nucleus as l increases.

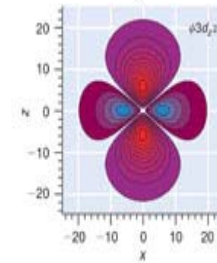
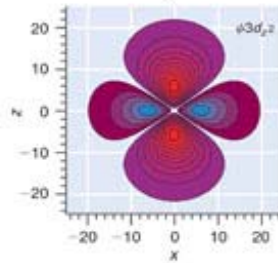
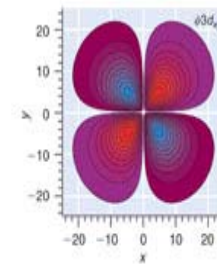
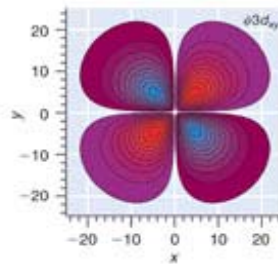
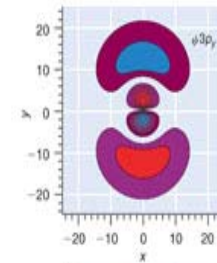
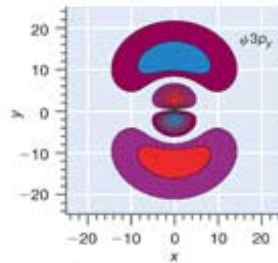
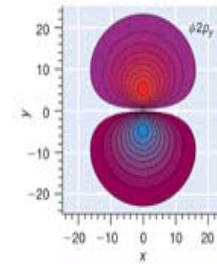
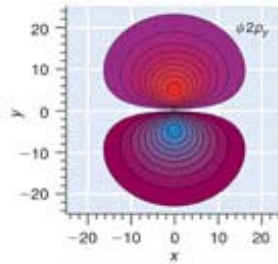
Note that the s orbital has a finite, non-zero value at the nucleus.

3.1-f: Contour Plots of the



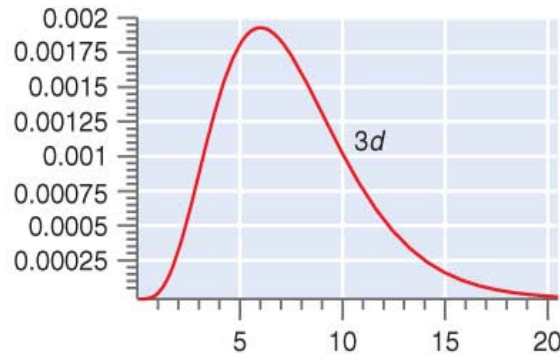
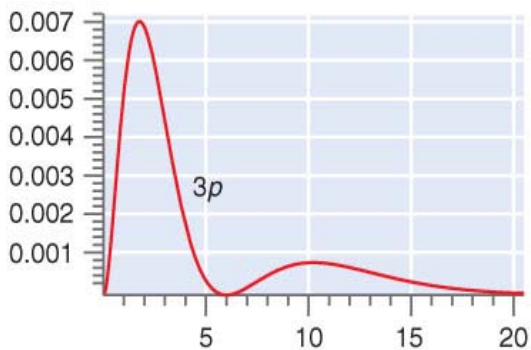
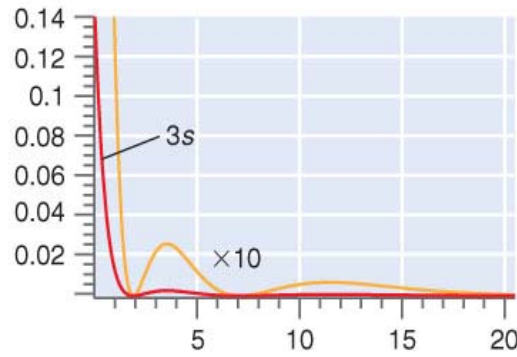
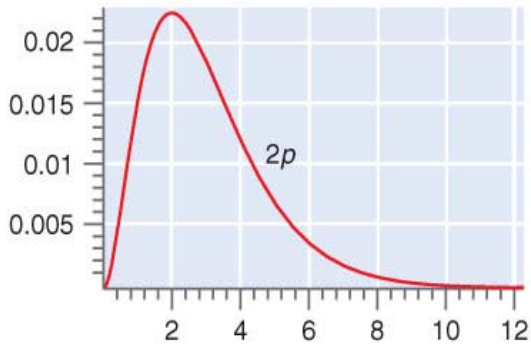
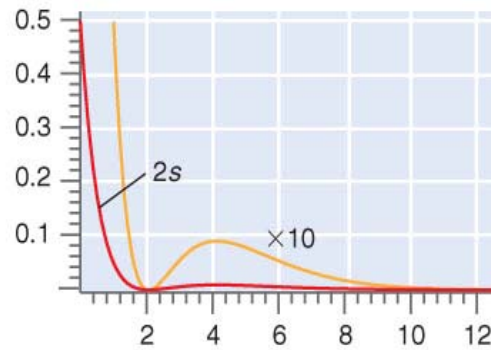
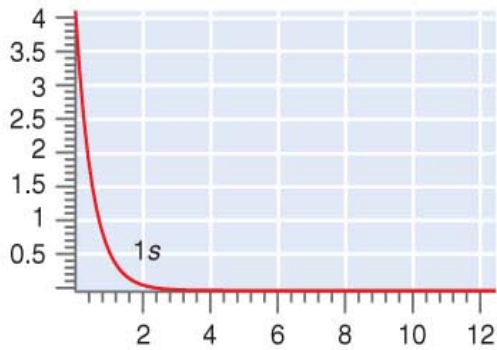
(b)

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3.1-g: Probability Density Functions

$$\Psi_{nlm_\ell}^2(r, \theta, \varphi)$$



- Shell model (the Bohr Atom) vs. Quantum mechanical model
Knowledge of the location of the electron in the H-atom vs. knowledge of the probability of finding it in a small volume element at a specific location

- The probability $\propto \Psi^* \Psi d\tau$

- To what extent does the exact quantum mechanical solution resemble the shell model?

⇒ The radial distribution function

3.1-h: The Radial Distribution Function

- The probability of finding the electron in a particular region, r , θ and φ , in space

$$\Psi_{n\ell m_\ell}^2(r, \theta, \varphi) \cdot r^2 \sin\theta \cdot dr \cdot d\theta \cdot d\varphi$$

- What we are really interested in is

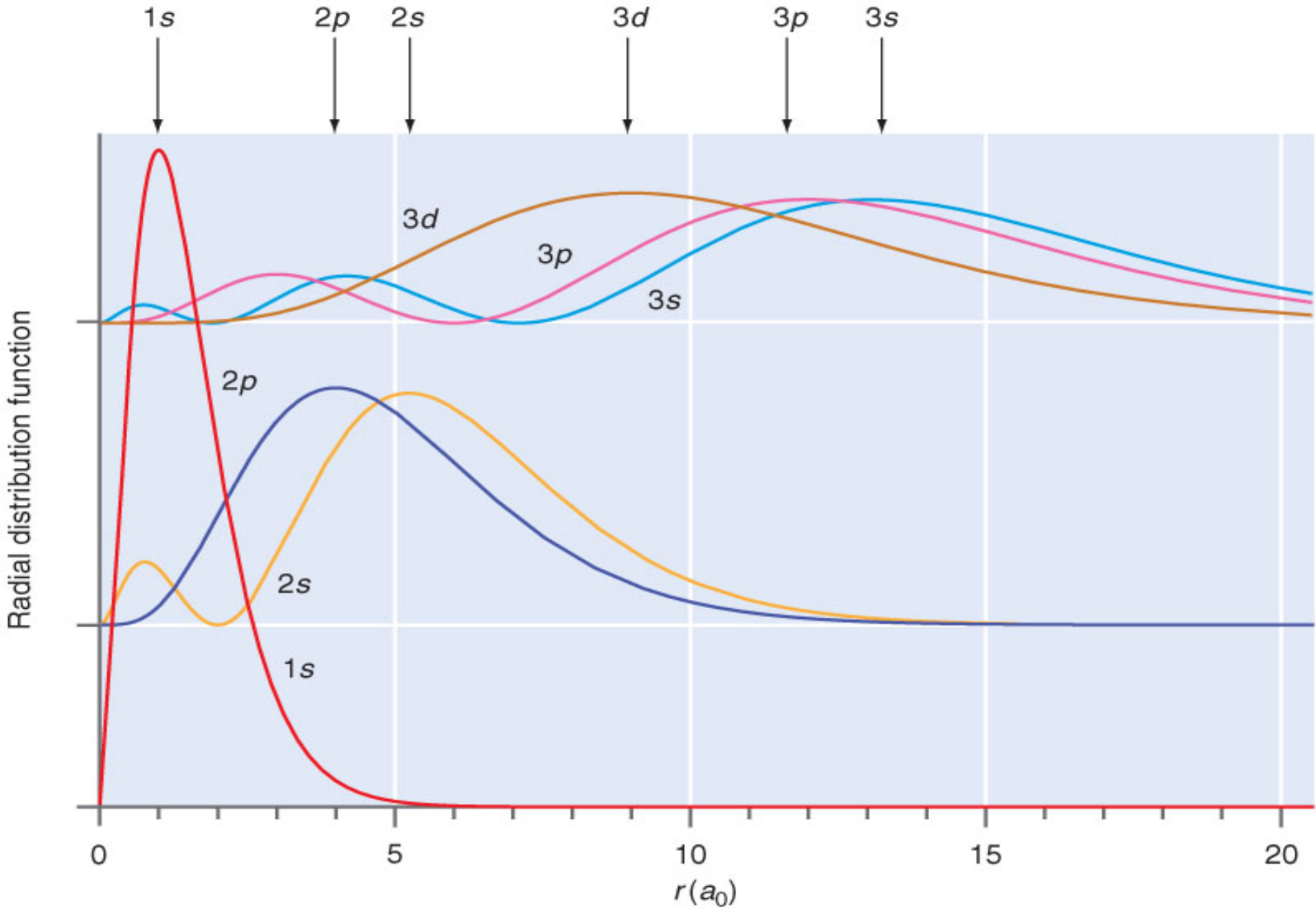
“What is the probability of finding the electron at a particular value of r , regardless of the values of θ and φ ?”

This can be obtained by integrating the probability density over all values of θ and φ

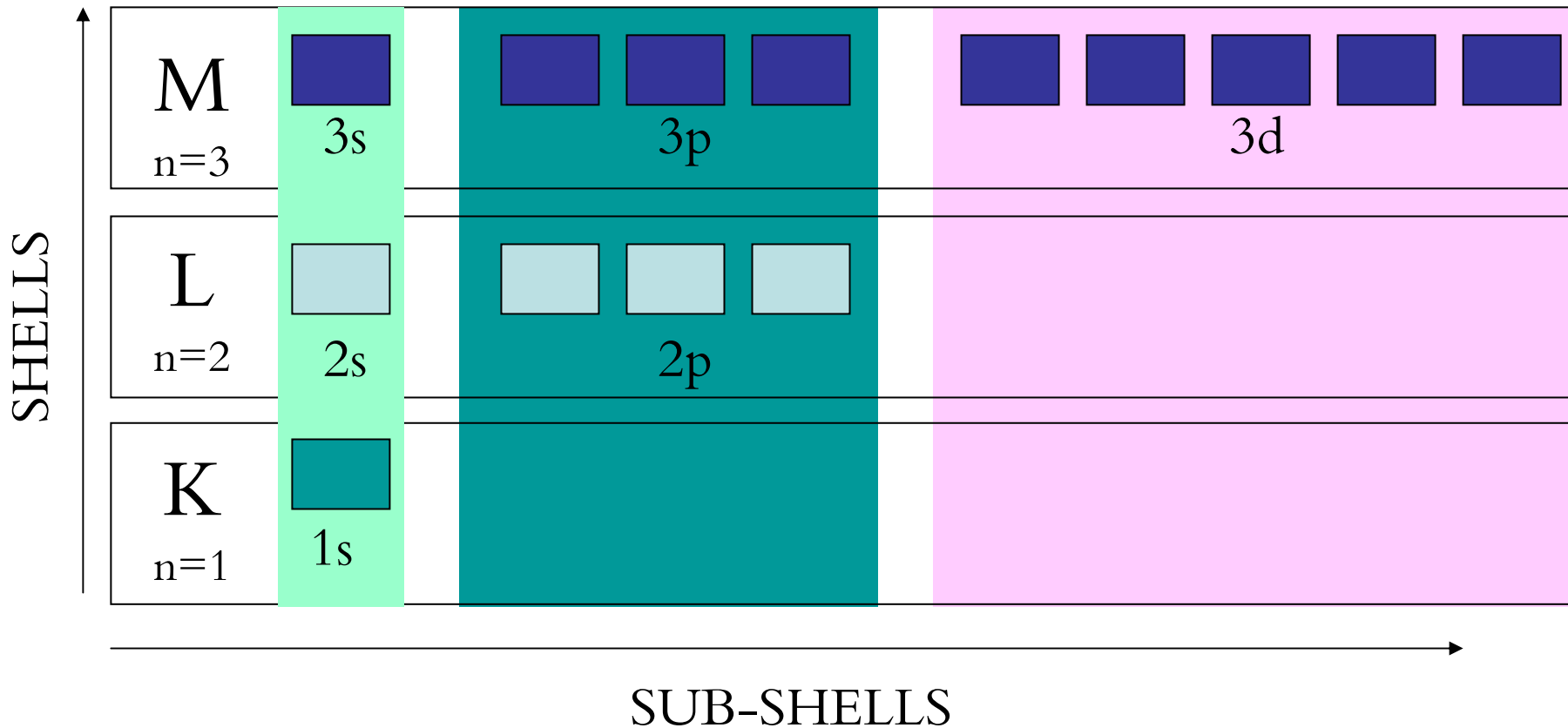
$$P(r) \cdot dr = \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} \Psi_{n\ell m_\ell}^2(r, \theta, \varphi) \cdot r^2 \sin\theta \cdot dr \cdot d\theta \cdot d\varphi$$
$$P(r) \cdot dr = r^2 \cdot [R_{n\ell}(r)]^2 \cdot dr$$

- The new function, $P(r)$, is called the **radial distribution function**

The probability function of choice to determine the most likely radius to find the electron for a given orbital



Old Convention: Shells vs. Sub-shells



SHELL

$n =$ 1 2 3 4
 \updownarrow \updownarrow \updownarrow \updownarrow
 K L M N

4.1 Angular Momentum

a) General Properties: apply to all angular momenta.

Any property \mathbf{M} is assigned as an angular momentum if it obeys the following relationships:

$$i) \hat{M}^2 = \hat{M}_x^2 + \hat{M}_y^2 + \hat{M}_z^2$$

$$ii) [\hat{M}_x, \hat{M}_y] = i\hbar\hat{M}_z; [\hat{M}_y, \hat{M}_z] = i\hbar\hat{M}_x; [\hat{M}_z, \hat{M}_x] = i\hbar\hat{M}_y$$

$$iii) [\hat{M}^2, \hat{M}_x] = [\hat{M}^2, \hat{M}_y] = [\hat{M}^2, \hat{M}_z] = 0$$

$$\mathbf{M} \text{ has } \hat{M}_+ = \hat{M}_x + i\hat{M}_y \quad \text{Raising operator}$$

$$\hat{M}_- = \hat{M}_x - i\hat{M}_y \quad \text{Lowering operator}$$

Two quantum numbers are needed because \mathbf{M} has two properties:

$$\hat{M}^2\psi_{j,m_j} = j(j+1)\hbar^2\psi_{j,m_j} \quad j = 0, \frac{1}{2}, 1, \frac{3}{2}, 2, \dots$$

$$\hat{M}_z\psi_{j,m_j} = m_j\hbar\psi_{j,m_j} \quad m_j = -j, -j+1, \dots, j-1, j$$

Note:

- 1.) j can have non-integer values in general
- 2.) m_j has symmetric values about 0

Also:

$$\langle \psi_{j,m_j+1} | \hat{M}_+ | \psi_{j,m_j} \rangle = \hbar \sqrt{j(j+1) - m_j(m_j+1)} = \hbar \sqrt{(j-m_j)(j+m_j+1)}$$

$$\langle \psi_{j,m_j-1} | \hat{M}_- | \psi_{j,m_j} \rangle = \hbar \sqrt{j(j+1) - m_j(m_j-1)} = \hbar \sqrt{(j+m_j)(j-m_j+1)}$$

Note: from the definitions of \mathbf{M}_+ and \mathbf{M}_- one can readily show:

$$\hat{M}_x = \frac{(\hat{M}_+ + \hat{M}_-)}{2} \quad \text{and} \quad \hat{M}_y = \frac{(\hat{M}_+ - \hat{M}_-)}{2i}$$

One specific example seen in C374a is Orbital Angular Momentum of the electron.

Here: $\hat{M} = \hat{L}$ and its operators on H-atom wave function ψ_{n,ℓ,m_ℓ}

The quantum numbers here are integer:

$$\ell = 0, 1, 2, \dots \quad m_\ell = -\ell, \dots, +\ell$$